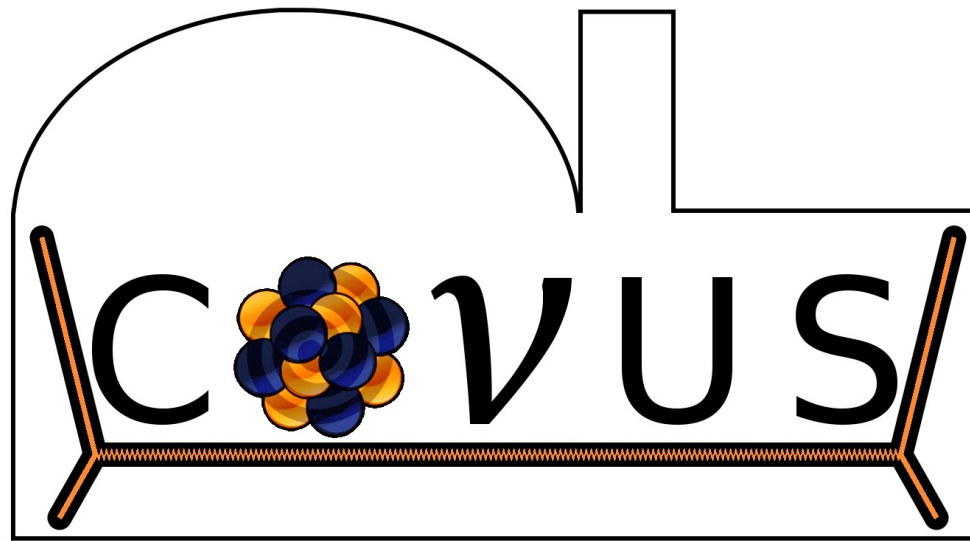




Coherent elastic neutrino-nucleus scattering

First constraints / observations and future potential

Thomas Rink (on behalf of the CONUS Collaboration)

Max-Planck-Institut für Kernphysik (MPIK), Heidelberg

56th Rencontres de Moriond 2022
12 to 19 March 2022



Coherent elastic neutrino-nucleus scattering (CEvNS)

CEvNS chronology:

- Daniel Freedman (1974): weak NC, flavor-blind, threshold-free!
- First detection with π -DAR: COHERENT CsI (2017) & LAr (2021)
- Reactor experiments (2019 - ...):
CONNIE (Si), CONUS (Ge), NCC-1701 (Ge)
- Several running and future experiments:
CCM (Ar), Miver (Ge/Si), NEON (NaI[Tl]), ν -cleus (CaWO₄, Al₂O₃),
 ν GEN (Ge), RED100 (Xe), Ricochet (Ge/Zn), Texono (Ge)

The Channel:

- Coherence = enhancement $\sim N^2$ $E_\nu \leq \frac{1}{2R_A} \approx \frac{197}{2.5\sqrt[3]{A}} [\text{MeV}]$
- Observable = nuclear recoil energy T_A

$$\frac{d\sigma}{dT_A} = \frac{G_F^2 m_A}{4\pi} \left[(1 - 4\sin^2 \theta_W)Z - N \right]^2 \left(1 - \frac{m_A T_A}{2E_\nu^2} \right) F^2(T_A)$$

$\rightarrow$ detector material

$\rightarrow$ neutrino source

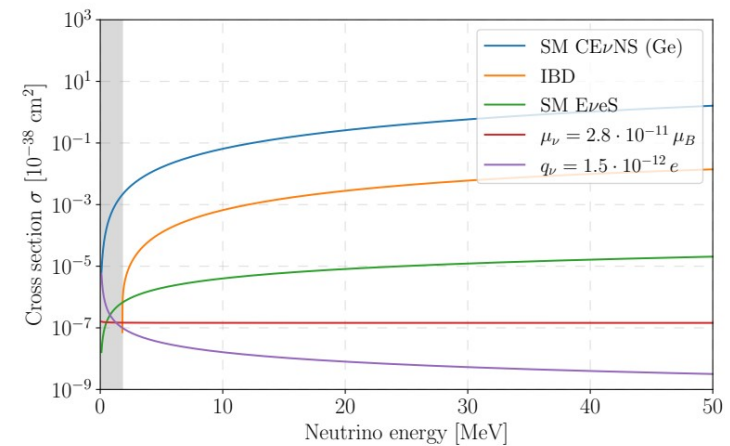
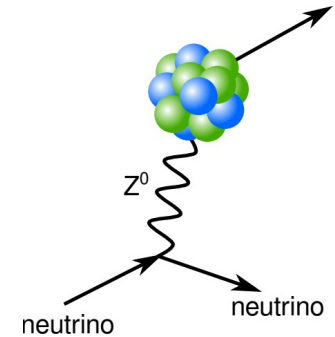
$\rightarrow$ energy threshold / neutrino source

- Very low energy threshold needed:

- $T_A \sim N^{-1}$

- Quenching: $T_A \rightarrow$ “detection”

Cross section σ
vs.
nuclear recoil T_A



[TR, PhD thesis, 2022]

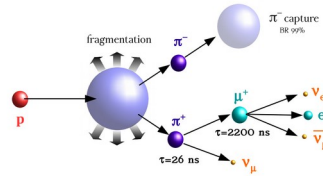
Element	N	E_ν [MeV]	T_A [keV]
Si	14	25.6	51.3
Ge	38 / 40 / 42	18.9	10.5
Xe	75 / 77 / 78	15.5	3.9
I / Cs	74 / 78	15.7 / 15.4	4.2 / 3.9

Two complementary paths

π -decay-at-rest neutrinos:

- Pulsed GeV-proton beam hitting heavy target
→ multiple ν flavors
- Time correlation of events
→ background suppression $\times (10^3\text{-}10^4)$
- Higher ν energies
→ larger cross section, but reduced coherence

→ **COHERENT, CCM, ESS...**

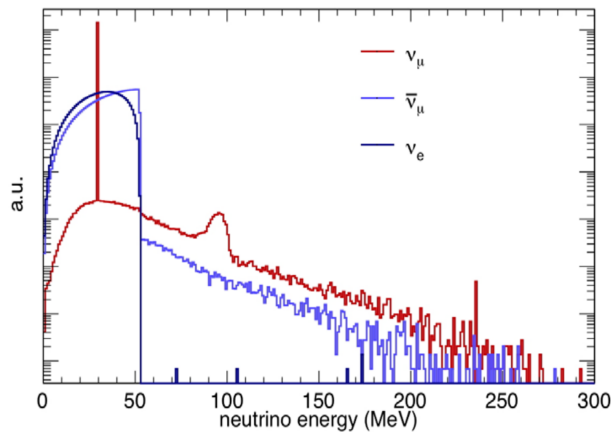


Reactor antineutrinos:

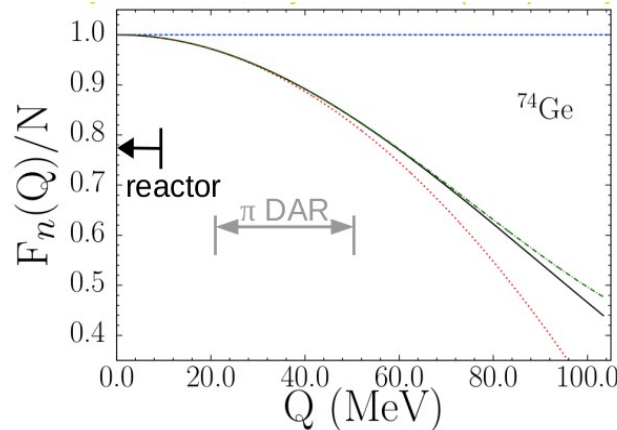
- β decays in nuclear reaction chains → only $\bar{\nu}_e$
- Strongest artificial ν source on earth:
 $\sim 10^{20}$ $\bar{\nu}_e$'s/GW/s
- ν energies up to 10 MeV → coherent regime!
- Close to reactor core: no lab conditions!
→ no cryogenic liquids, no remote control

→ **CONNIE, CONUS, NCC-1701, ...**

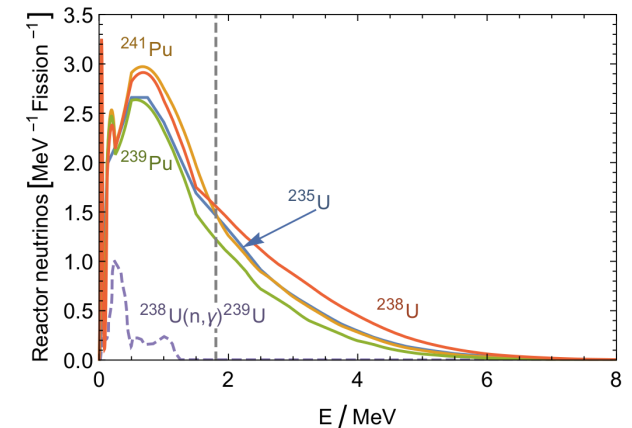
[Akimov et al., [arXiv:1509.08702](https://arxiv.org/abs/1509.08702), 2015]



[Patton et al., [10.1103/PhysRevC.86.024612](https://arxiv.org/abs/10.1103/PhysRevC.86.024612), 2012]



[Vitagliano et al., [10.1103/RevModPhys.92.045006](https://arxiv.org/abs/10.1103/RevModPhys.92.045006), 2020]



Beam-reactor complementarity:

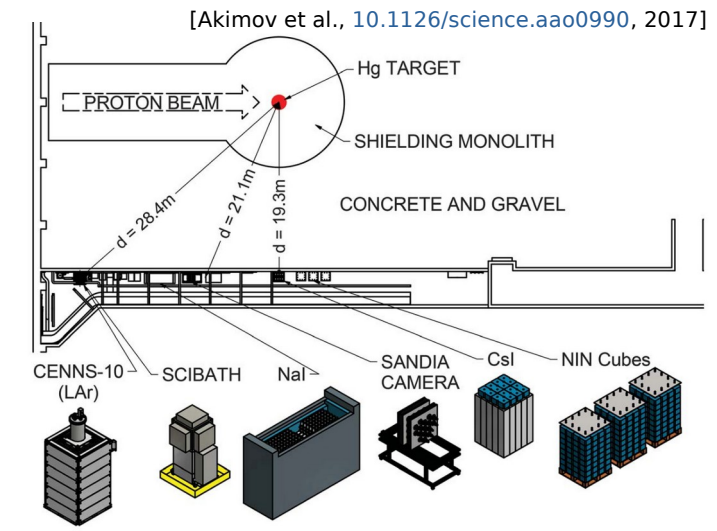
CEvNS at reactor site as high statistic baseline for multi-target and multi-flavored beam investigations!

Experimental efforts

COHERENT “neutrino alley” at Oak Ridge National Lab (USA)

- Overburden: 8m w.e.
- ν flux: $4.3 \cdot 10^7/\text{cm}^2/\text{s}$ @ 20m
- Multiple targets @ different baselines

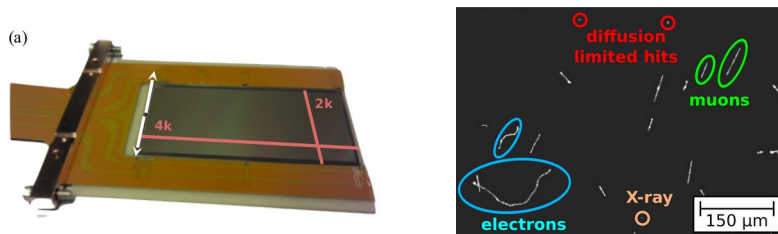
Target	N	Mass [kg]	$T^{\text{thr}} [\text{keV}_{\text{nr}}]$
CsI[Na]	78/74	14.6	~ 6.5
LAr	22	22	~ 20
Nal[Tl]	12/74	185	~ 13
Ge	38/40/42	10	~ 5



CONNIE (Brazil)

- CCDs (Si, $N=14$) : 47.6g, $T_A^{\text{thr}} \sim 1 \text{keV}_{\text{nr}}$
- @30m to 3.8GW_{th} PWR:
 ν flux: $7.8 \cdot 10^{12}/\text{cm}^2/\text{s}$

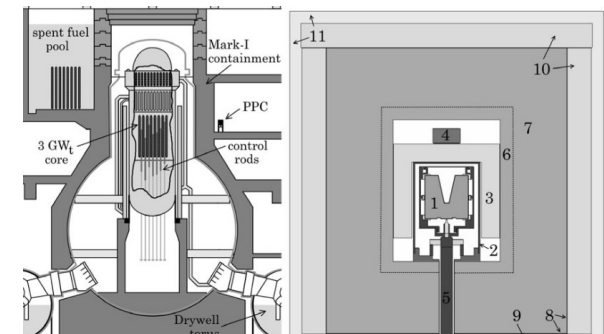
[Aguilar-Arevalo et al., 10.1088/1742-6596/761/1/012057, 2016]



[Aguilar-Arevalo et al., 10.1088/1748-0221/11/07/P07024, 2016]

NCC-1701 (USA)

- Ge PPC ($N=\{38,40,42\}$):
3kg, $T_A^{\text{thr}} \sim 1\text{-}2 \text{keV}_{\text{nr}}$
- @10.3m to 2.96GW_{th} BWR:
 ν flux: $5 \cdot 10^{13}/\text{cm}^2/\text{s}$



[Colaesi et al., 10.1103/PhysRevD.104.072003, 2021]

The CONUS experiment

[Bonet et al., arXiv:2110.02174, 2021]

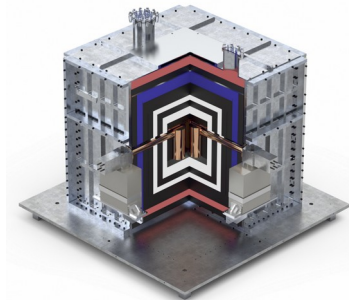
Reactor: Brokdorf nuclear power plant

- @17m to 3.9GW_{th}: inside reactor building!
- ν flux: $2.3 \cdot 10^{13}/\text{cm}^2/\text{s}$
- Reactor information:
fuel composition, thermal power

Background & shield

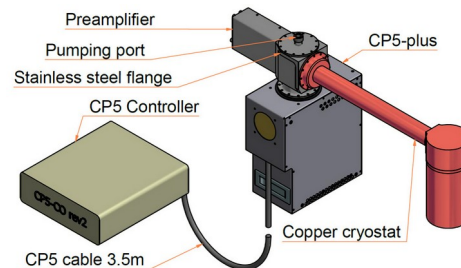
[Bonet et al.,
arXiv:2112.09585, 2021]

- Overburden: 10-45 m w.e.
- Active and passive components:
bkg reduction $\times 10^3$ - 10^4
- Critical reactor-correlated bkg
under control!
- Background level $O(10)/\text{keV}/\text{d}/\text{kg}$

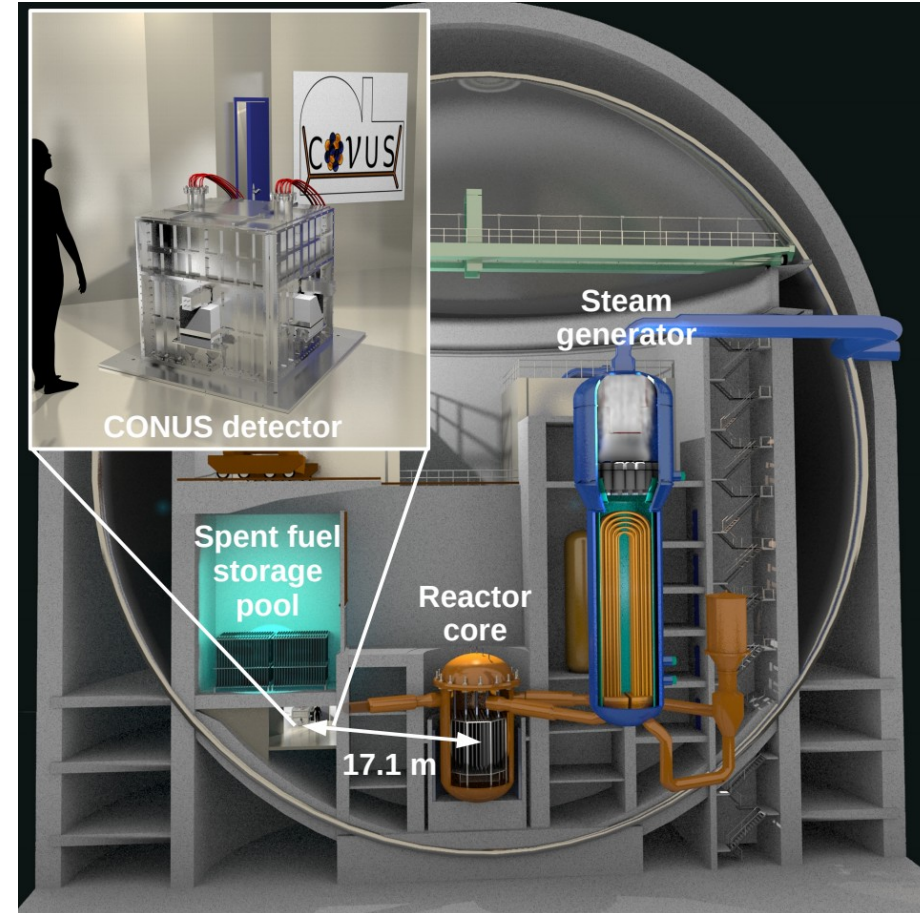


Detectors:

- HPGe PPC: 4*1kg
- $E_{\text{thr}} \sim 300\text{eV} \rightarrow T_A^{\text{thr}} \sim 2\text{keV}_{\text{nr}}$
- Low Bkg design
- Electrical cryocooler



[Bonet et al.,
10.1140/epjc/s10052-021-09038-3,
2021]

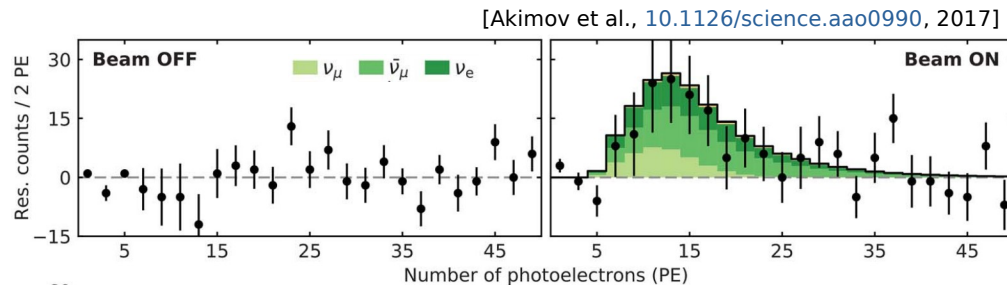


Close to reactor core:
lab conditions not easily realized
→ site monitoring crucial!

CEvNS observations and constraints

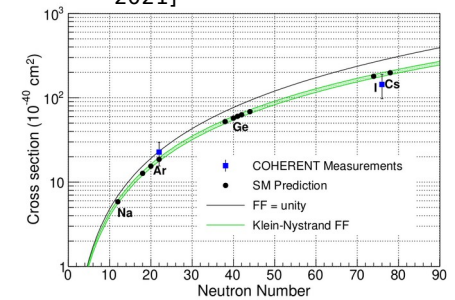
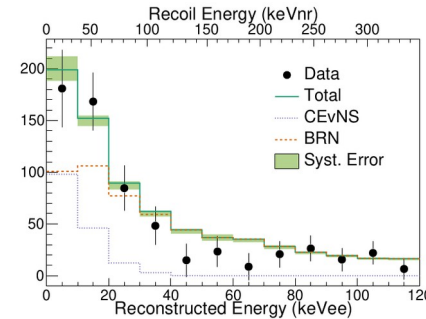
Observations:

COHERENT CsI[Na] (2017): $5 * 10^{20}$ POT
 134 ± 22 (pred. 173 ± 48)



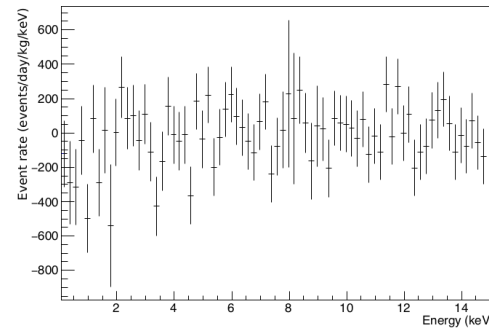
COHERENT LAr (2021): $13.7 * 10^{22}$ POT
 159 ± 43 (pred. 128 ± 17)

[Akimov et al.,
[10.1103/PhysRevLett.126.012002](https://doi.org/10.1103/PhysRevLett.126.012002),
 2021]

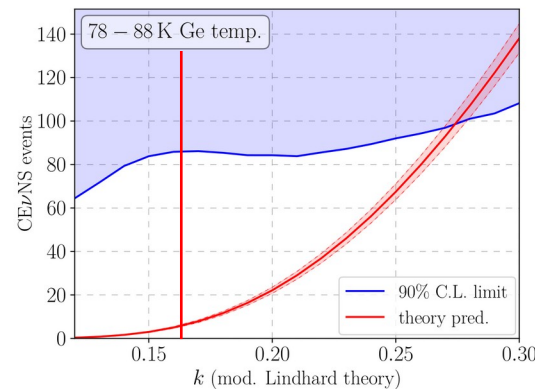


Limits:

CONNIE Si (2019):
 $R_{NP} < 40 * R_{SM}$ @ 95% C.L.
 2.1kg*d ON + 1.6kg*d OFF
 [Aguilar-Arevalo et al.,
[10.1103/PhysRevD.100.092005](https://doi.org/10.1103/PhysRevD.100.092005), 2019]

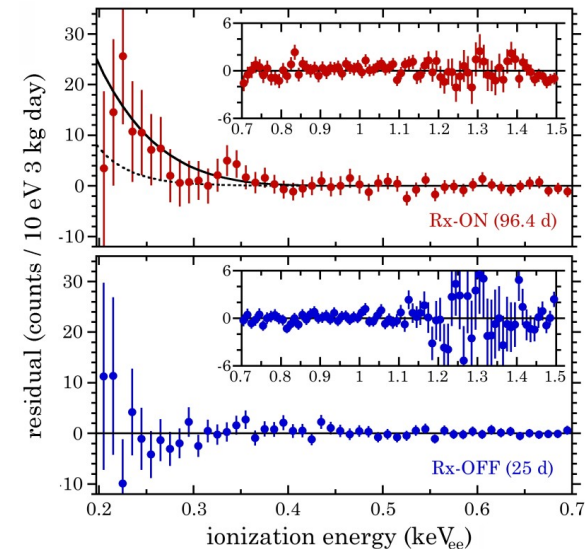


CONUS Ge (2020):
 $R_{SM} < 0.4/\text{kg/d}$ @ 90% C.L.
 (x17 lower then prediction)
 249kg*d ON + 59kg*d OFF
 [Bonet et al.,
[10.1103/PhysRevLett.126.041804](https://doi.org/10.1103/PhysRevLett.126.041804), 2020]



“Suggestive evidence for Coherent Elastic Neutrino-Nucleus Scattering from reactor antineutrinos” [Colaresi et al., [arXiv:2202.09672](https://arxiv.org/abs/2202.09672), 2022]

NCC-1701 Ge:
 289kg*d ON
 +
 75kg*d OFF



Non-standard interactions (NSIs)

Aim: Test for new (high-scale) neutrino interaction

[Barranco et al., [10.1088/1126-6708/2005/12/021](https://doi.org/10.1088/1126-6708/2005/12/021), 2005]

- Vector-type Interaction $\mathcal{O}_{\alpha\beta}^{qV} = (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{q} \gamma_\mu P_{L/R} q) + \text{h.c.}$
→ New feature: destructive interference!
- Modified of weak nuclear charge
$$\mathcal{Q}_{NSI}^V = (2\epsilon_{\alpha\alpha}^{uV} + \epsilon_{\alpha\alpha}^{dV} + g_p^V) Z + (\epsilon_{\alpha\alpha}^{uV} + 2\epsilon_{\alpha\alpha}^{dV} + g_n^V) N$$

$$+ \sum_{\alpha \neq \beta} \left[(2\epsilon_{\alpha\beta}^{uV} + \epsilon_{\alpha\beta}^{dV}) Z + (\epsilon_{\alpha\beta}^{uV} + 2\epsilon_{\alpha\beta}^{dV}) N \right]$$
- New couplings ϵ partially degenerate
→ different isotopes/targets at different sources

$$\Lambda_{NP} > 100 \text{ GeV} \quad (\text{CONUS, } k=0.16)$$

$$\Lambda_{NP} > 240 \text{ GeV} \quad (\text{COHERENT, LAr})$$

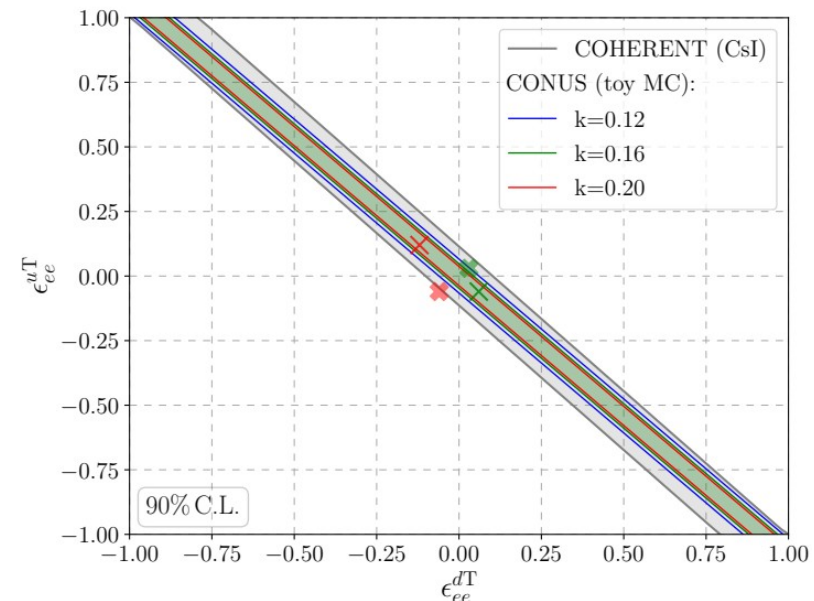
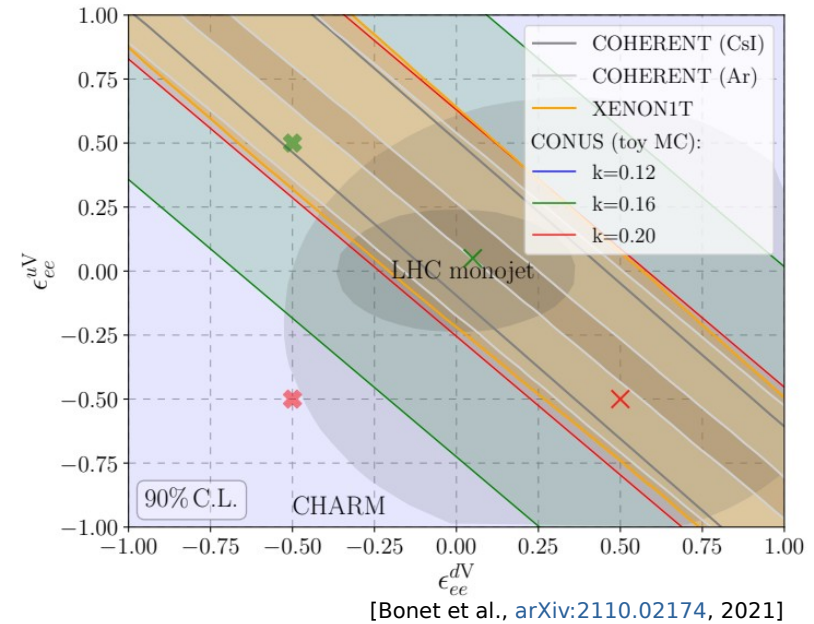
- Tensor-type operator $\mathcal{O}_{\alpha\beta}^{qT} = (\bar{\nu}_\alpha \sigma^{\mu\nu} \nu_\beta) (\bar{q} \sigma_{\mu\nu} q) + \text{h.c.}$
[Barranco et al., [10.1142/S0217751X12501473](https://doi.org/10.1142/S0217751X12501473), 2012]
- Interaction with higher kinematic cut-off
$$\left(\frac{d\sigma}{dT_N} \right)_{NSI}^T = \frac{G_F^2 M}{\pi} (\mathcal{Q}_{NSI}^T)^2 \left(1 - \frac{MT_N}{4E_\nu^2} \right)$$

→ quenching less relevant, overall background crucial

$$\Lambda_{NP} > 360 \text{ GeV} \quad (\text{CONUS, } k=0.16)$$

If sub-percent sensitivity on $\epsilon \rightarrow$ probing **TeV-scale** physics!

[Lindner et al., [10.1007/JHEP03\(2017\)097](https://doi.org/10.1007/JHEP03(2017)097), 2017]



Simplified models: light mediators at low-E

Aim: Test specific but simple mediator models that contribute to CEvNS

[Cerdeño et al.,
10.1007/JHEP05(2016)118, 2016]

Light scalar:

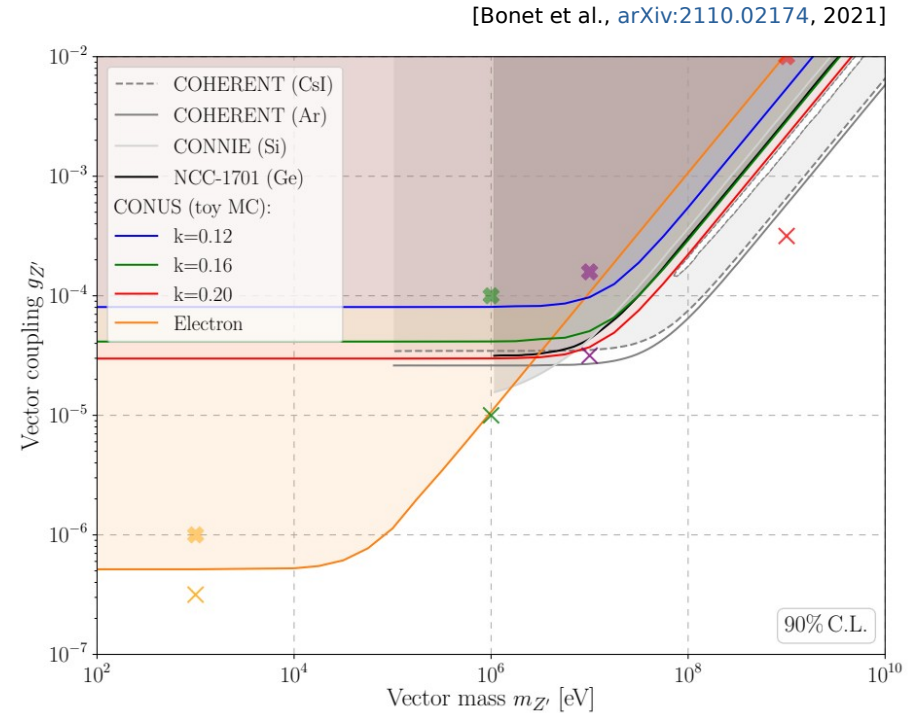
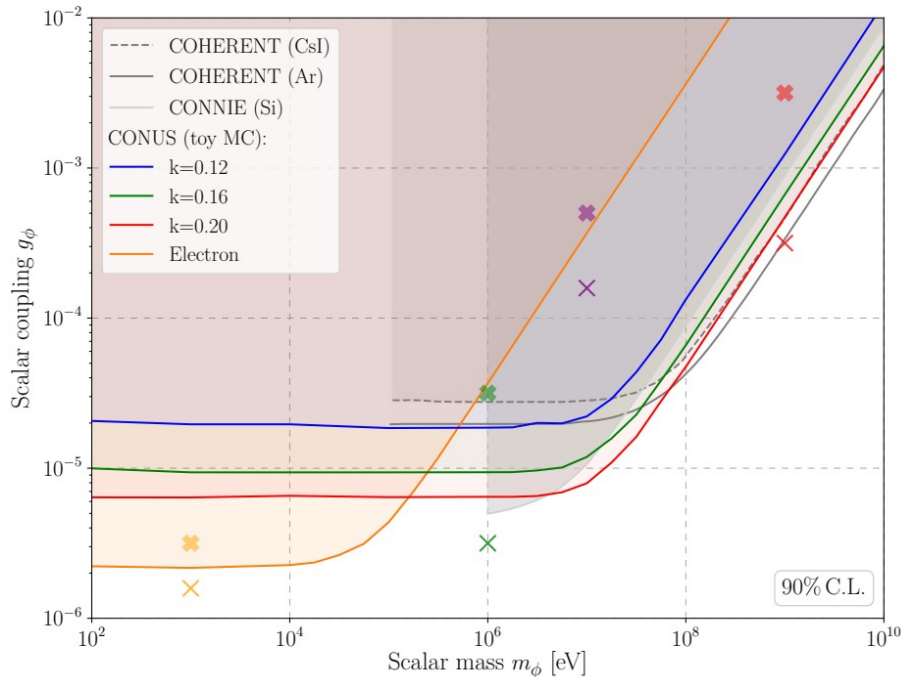
$$\mathcal{L}_\phi = \phi \left(g_\phi^{qS} \bar{q}q + g_\phi^{eS} \bar{e}e + g_\phi^{\nu S} \bar{\nu}_R \nu_L + \text{h.c.} \right) - \frac{1}{2} m_\phi^2 \phi^2$$

- Assume: universal coupling to quarks / neutrinos:
- spectral distortions for **small recoil energies**

Light vector:

$$\mathcal{L}_{Z'} = Z'_\mu \left(g_{Z'}^{\nu V} \bar{\nu}_L \gamma^\mu \nu_L + g_{Z'}^{eV} \bar{e} \gamma^\mu e + g_{Z'}^{qV} \bar{q} \gamma^\mu q \right) + \frac{1}{2} m_{Z'}^2 Z'_\mu Z'^\mu$$

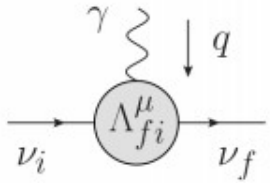
$(m_\phi, g_\phi), (m_{Z'}, g_{Z'})$



Reactor neutrinos for low masses $\leftrightarrow$ π -DAR neutrinos for higher masses

Electromagnetic neutrino properties

Loop-induced electromagnetic properties: [Giunti & Studenikin, [10.1103/RevModPhys.87.531](https://arxiv.org/abs/10.1103/RevModPhys.87.531), 2015]



- neutrino charge (radius), electric and magnetic dipole moment, anapole moment
- characteristic dependence on neutrino's fermionic nature
- test BSM physics!

ν MM from electron scattering at reactor site:

$$\left(\frac{d\sigma}{dT_e}\right)_{\nu\text{MM}} = \frac{\pi\alpha_{\text{em}}^2}{m_e^2} \left(\frac{1}{T_e} - \frac{1}{E_\nu}\right) \left(\frac{\mu_{\nu e}}{\mu_B}\right)^2$$

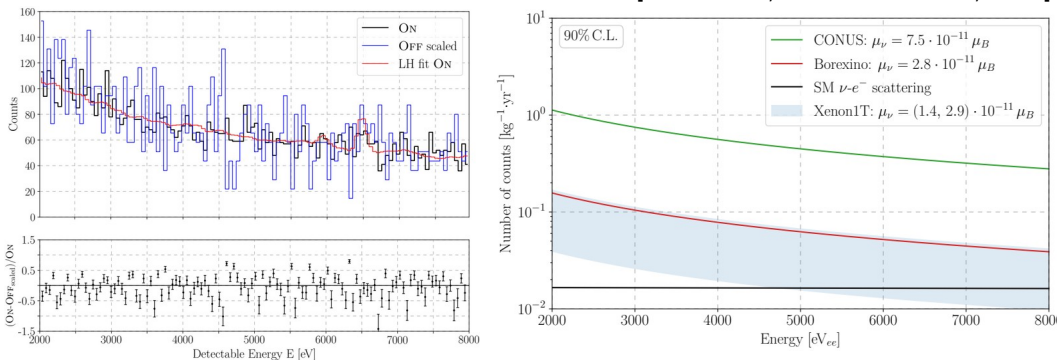
- huge reactor ν flux
- low threshold energies!

ν MM from CEvNS at a π -DAR site:

$$\left(\frac{d\sigma}{dT_A}\right)_{\nu\text{MM}} = \frac{\pi\alpha_{\text{em}}^2}{m_e^2} Z^2 \left(\frac{1}{T_A} - \frac{1}{E_\nu}\right) \left(\frac{\mu_{\nu e}}{\mu_B}\right)^2 F^2(T_A)$$

- higher ν energy
- multiple flavors

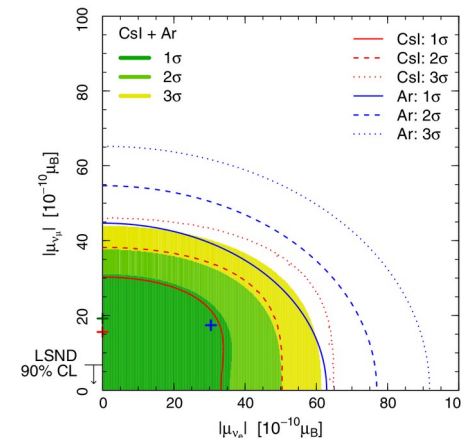
[Bonet et al., [arXiv:2201.12257](https://arxiv.org/abs/2201.12257), 2022]



$$\mu_{\nu e} < 7.5 \cdot 10^{-11} \mu_B \text{ (90\% C.L.)}$$

Assuming null result: convert to ν MC bound

$$|q_{\nu e}| < 3.3 \cdot 10^{-12} e \text{ (90\% C.L.)}$$



[Cadeddu et al., [10.1103/PhysRevD.102.015030](https://arxiv.org/abs/10.1103/PhysRevD.102.015030), 2020]

ν MC bounds
from COHERENT:

$$|q_\nu| \sim \mathcal{O}(10^{-8})e$$

Future studies of transition moments,
neutrino charge radii..., etc.

Opportunities for CEvNS experiments

SM investigations:

- **Weinberg angle at low-Q:**

Test for deviations at low energy
→ light Z' , neutrino charge radius

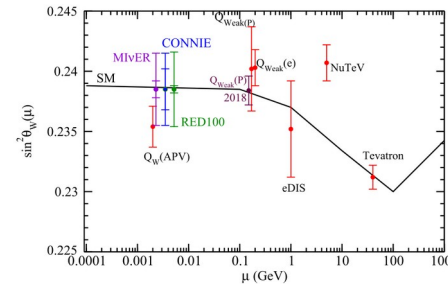
- **Nuclear form factors:**

model-independent extraction of neutron density distributions
→ Beam-reactor complementarity

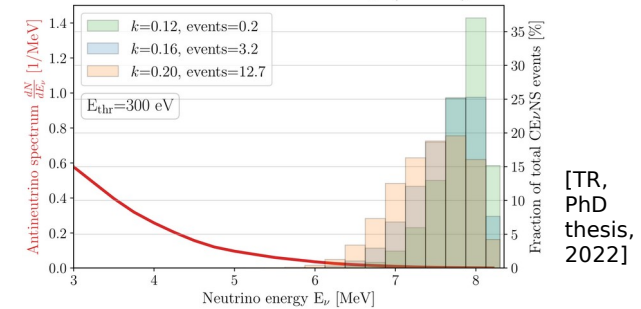
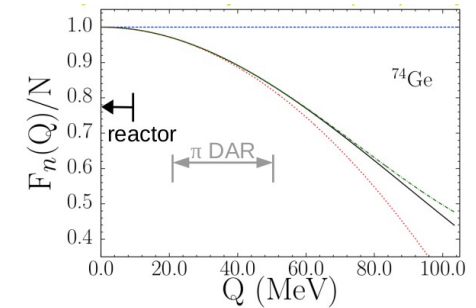
- **Measuring reactor antineutrino spectrum:**

CEvNS sensitive to high-E part where uncertainty is largest

[Cañas et al., 10.1016/j.physletb.2018.07.049, 2018]



[Patton et al., 10.1103/PhysRevC.86.024612, 2012]



[TR, PhD thesis, 2022]

BSM investigations:

- **Light sterile neutrinos:**

Use CEvNS for ν flux measurements

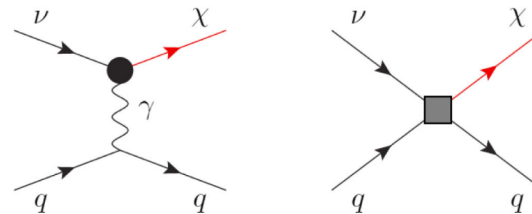
- **New fermion searches:**

Test further ν interactions
→ ν mass, DM, ...

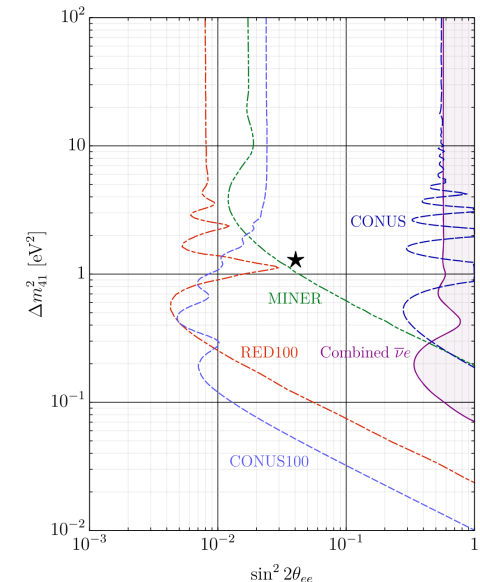
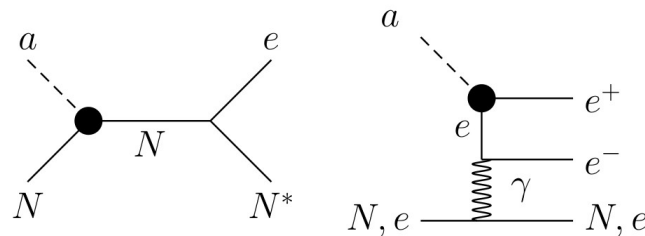
- **Probing portals:**

ALPs, dark photons, etc.

[Chang & Liao, 10.1103/PhysRevD.102.075004, 2022]



[Aristizabal Sierra et al., 10.1007/JHEP03(2021)294, 2021]



[Berryman, 10.1103/PhysRevD.100.023540, 2019]

Conclusion

- **CEvNS opens new path to high-statistics neutrino physics**

- beams and reactors complement each other!
- full spectrum of modern detection technologies
- “car-size” neutrino detectors!

- **CEvNS data allow various SM & BSM investigations**

- Weinberg angle, nuclear form factors, ...
- NSIs, light mediators, neutrino properties, DM, ...

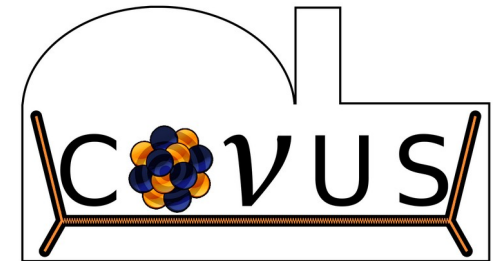
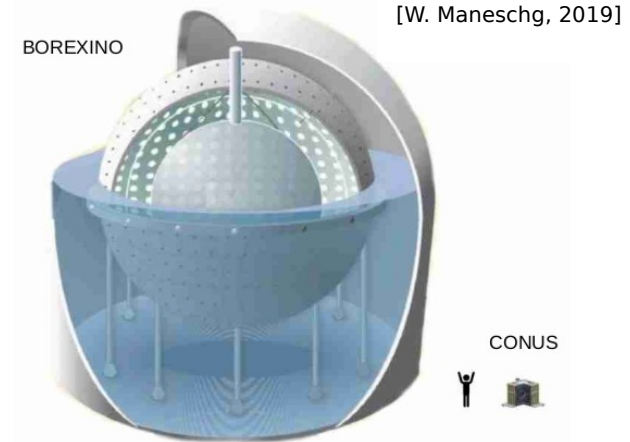
- **Several improvements of the CONUS set-up:**

- environmental stability, new DAQ (lower E_{thr} , pulse shapes)
- analysis of new data in preparation

→ new experimental site under discussion!

- **Future: flavor-blind neutrino astronomy**

- supernovae, neutrino floor, ...



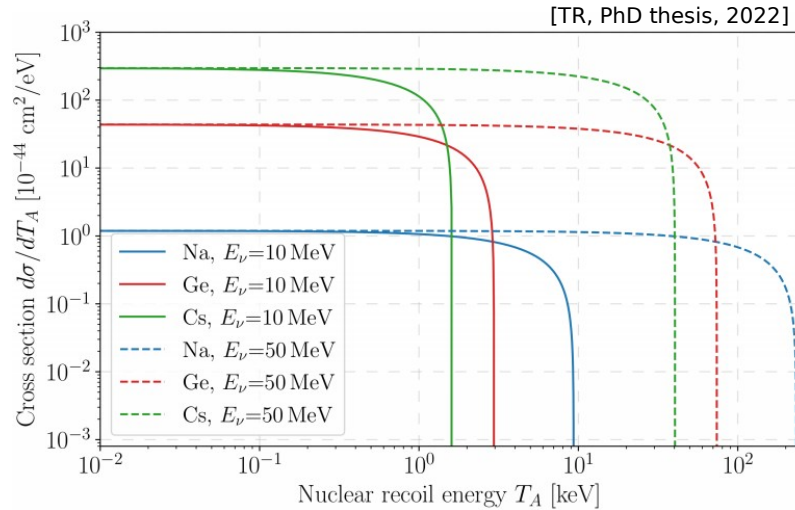
[Hakenmüller et al. [10.1140/epjc/s10052-019-7160-2](https://arxiv.org/abs/10.1140/epjc/s10052-019-7160-2), 2019]
[Bonet et al., [10.1103/PhysRevLett.126.041804](https://arxiv.org/abs/10.1103/PhysRevLett.126.041804), 2020]
[Bonet et al., [10.1140/epjc/s10052-021-09038-3](https://arxiv.org/abs/10.1140/epjc/s10052-021-09038-3), 2021]
[Bonet et al., [arXiv:2110.02174](https://arxiv.org/abs/2110.02174), 2021]
[Bonet et al., [arXiv:2112.09585](https://arxiv.org/abs/2112.09585), 2021]
[Bonet et al., [arXiv:2201.12257](https://arxiv.org/abs/2201.12257), 2022]
[Bonhomme et al., [arXiv:2202.03754](https://arxiv.org/abs/2202.03754), 2022]

Next-generation experiments (π -DAR, reactors & DM DD) = huge playground for phenomenology

(Neutrino magnetic moment, neutrino charge radius, vector NSI, tensor NSI, light scalars, light vectors, sterile neutrinos, dark matter, weak mixing angle, nuclear form factors, neutrino-electron-scattering, supernova neutrinos, antineutrino reactor spectra, neutrino floor, etc.)

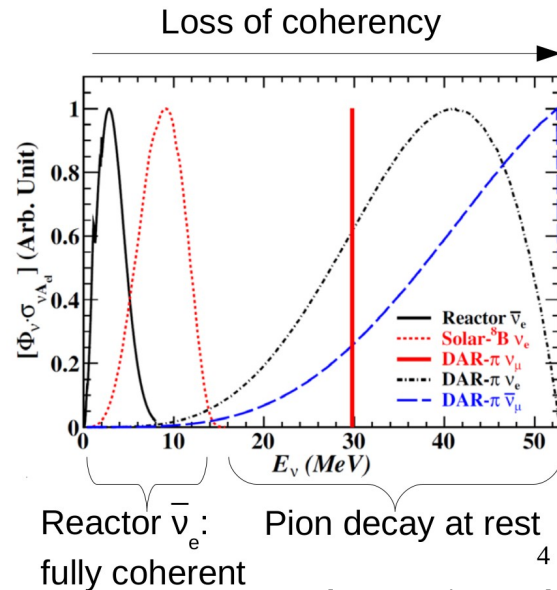
Backup

CEvNS with different sources and targets



Element	N	r_A [fm]	$E_\nu^{\max}$ [MeV]	$T_A^{\max}$ [keV]
Na	12	3.6	27.7	71.5
Si	14	3.8	25.9	51.3
Ar	22	4.4	23.1	28.5
Ge	38/40/42	5.2	18.9	10.5
I	74	6.3	15.7	4.16
Xe	75/77/78	6.4	15.5	3.93
Cs	78	6.4	15.4	3.85

[TR, PhD thesis, 2022]



[W. Maneschg, 2017]

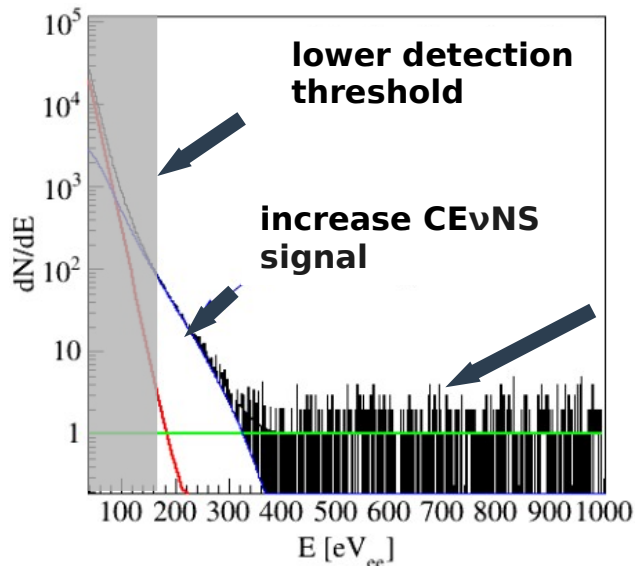
Neutrino source	Target	$T_A^{\max}$ [keV]	$E(QF \in \{0.1, 0.15, 0.2\})$ [keV]
Nuclear reactor (10 MeV)	Na	9.33	0.93 / 1.40 / 1.87
	Si	7.64	0.76 / 1.15 / 1.53
	Ar	5.37	0.54 / 0.81 / 1.07
	Ge	2.96	0.30 / 0.44 / 0.59
	I	1.69	0.17 / 0.25 / 0.34
	Xe	1.64	0.16 / 0.25 / 0.33
	Cs	1.62	0.16 / 0.24 / 0.32
π -DAR source (50 MeV)	Na	232.4	23.2 / 34.9 / 46.5
	Si	190.4	19.0 / 28.6 / 38.1
	Ar	134.0	13.4 / 20.1 / 26.8
	Ge	73.8	7.38 / 11.1 / 14.8
	I	42.3	4.23 / 6.34 / 8.45
	Xe	40.9	4.09 / 6.13 / 8.17
	Cs	40.4	4.04 / 6.03 / 8.07

Experimental requirements at reactor site

Goal: Detecting CEvNS with high accuracy!

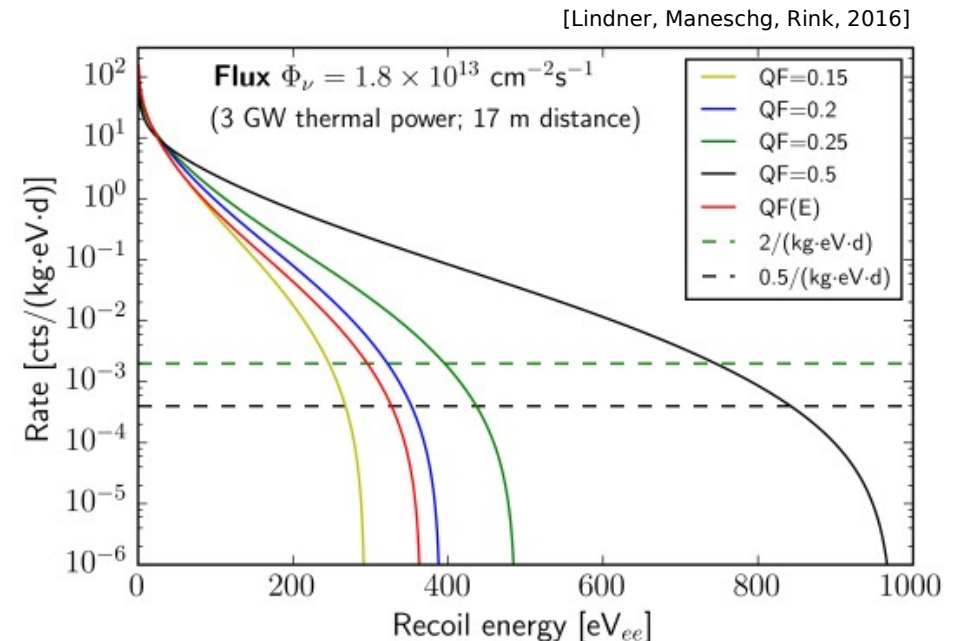
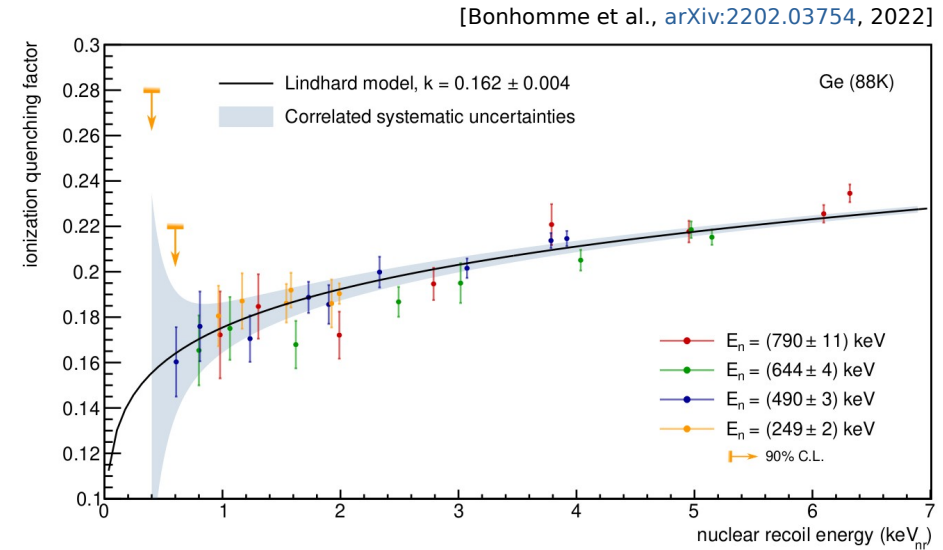
Several obstacles to overcome:

- 1) Beat $1/R^2$ factor
 - strong (= commercial) power plant, close to reactor core
- 2) Compensate quenching ($E_{\text{recoil}} \rightarrow E_{\text{ion}}$)
 - lowest possible detection threshold
- 3) Low background outside lab conditions
 - moderate overburden & limited shielding capacities



Reduce background level

E_ν : 10MeV
 $\max E_{\text{Recoil}}$: 3keV
 $\max E_{\text{ion}}$: ~600eV



Antineutrinos from nuclear reactions

Antineutrino emission in β decays of fuel reaction chain

- Mainly from ^{235}U , ^{238}U , ^{239}Pu , ^{241}Pu $\rightarrow >99\%$
- $\sim 6\text{-}7$ ν 's/fission up to 10 MeV
- Spectral distribution

$$S(E_\nu) = \frac{1}{4\pi R^2} \frac{W_{th}}{\sum_i \alpha_i E_i} \sum_i \alpha_i \left(\frac{dN_i}{dE_\nu} \right)$$

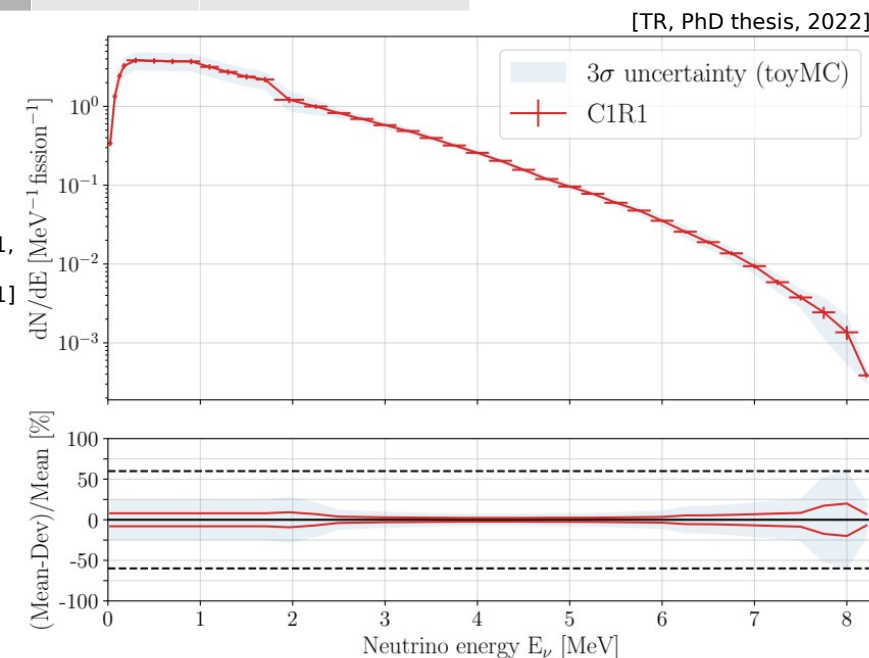
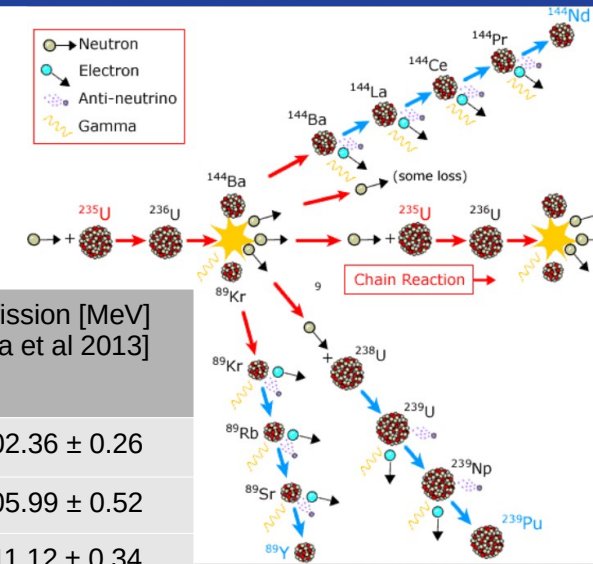
Knowledge about a reactors emission spectra

- Summation methods [Kopeikin et al. [10.1134/1.1825513](https://doi.org/10.1134/1.1825513), 2004]
 $\rightarrow$ summing β branches of all fission fragments
- Conversion methods [Haag et al., [10.1103/PhysRevLett.112.122501](https://doi.org/10.1103/PhysRevLett.112.122501), 2014
 Huber, [10.1103/PhysRevC.85.029901](https://doi.org/10.1103/PhysRevC.85.029901), 2011,
 Mueller et al., [10.1103/PhysRevC.83.054615](https://doi.org/10.1103/PhysRevC.83.054615), 2011]
 $\rightarrow$ measure β decay electron spectrum and convert into ν spectrum
- Direct measurements (IBD) [An et al., [10.1088/1674-1137/41/1/013002](https://doi.org/10.1088/1674-1137/41/1/013002), 2017]

Reality much more complicated...

- Varying reactor power $\rightarrow \mathbf{P(t)}$
- Changing fuel composition $\rightarrow \mathbf{a(t)}$

isotope	Fission fraction α (PWR)	E/fission [MeV] [Ma et al 2013]
^{235}U	57%	202.36 ± 0.26
^{238}U	8%	205.99 ± 0.52
^{239}Pu	30%	211.12 ± 0.34
^{241}Pu	5%	214.26 ± 0.33

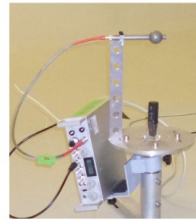
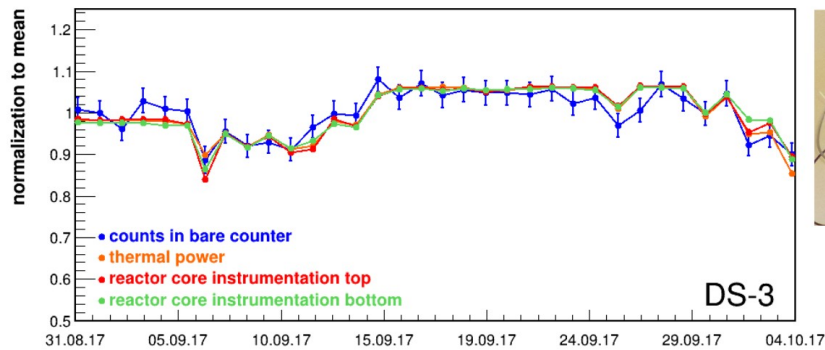


Power-correlated reactor radiation

→ Dedicated investigation of reactor-correlated background contributions:

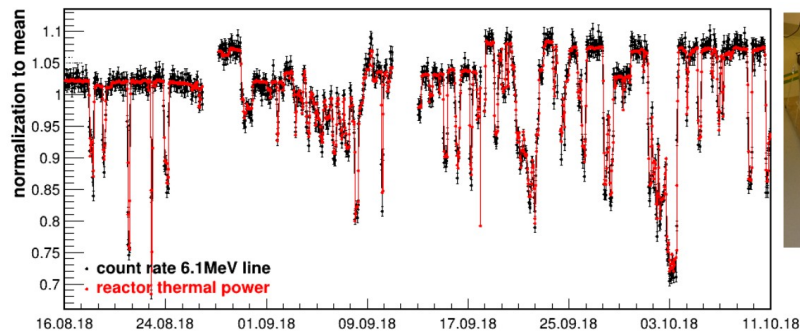
[Hakenmüller et al.
10.1140/epjc/s10052-019-7160-2, 2019]

- Simulation and validation of neutrons emitted from reactor core (at CONUS site) → thermal neutron counter

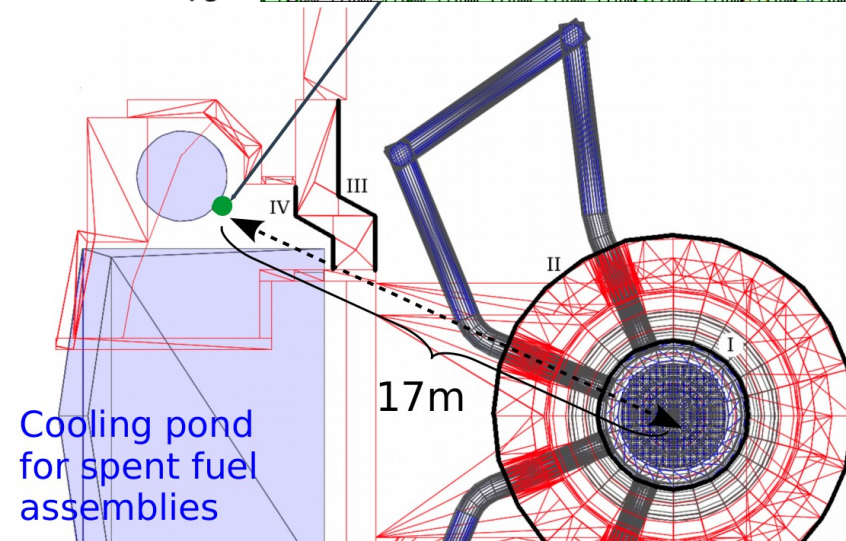
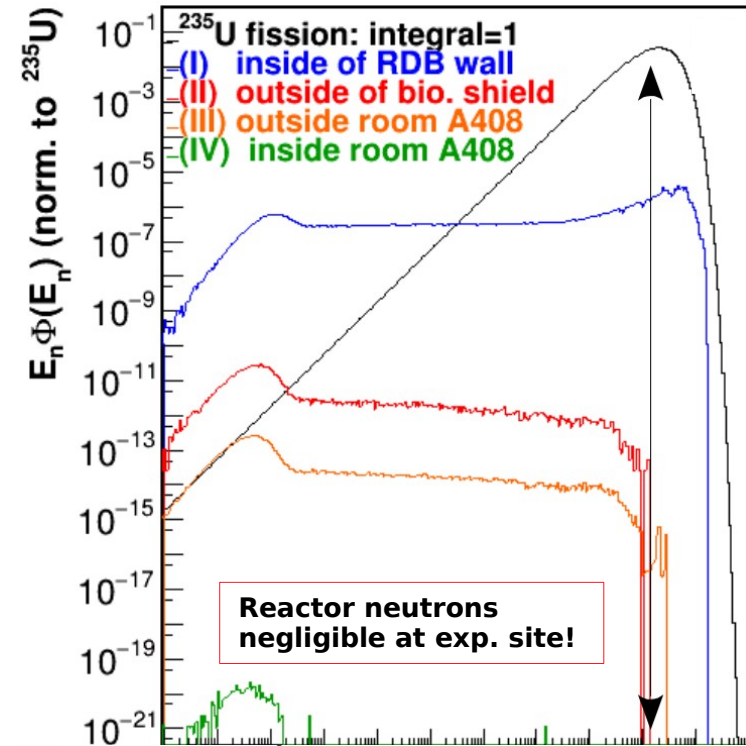


Correlation to
incore and excore
instrumentation

- γ radiation (6.1 MeV from ^{16}N)

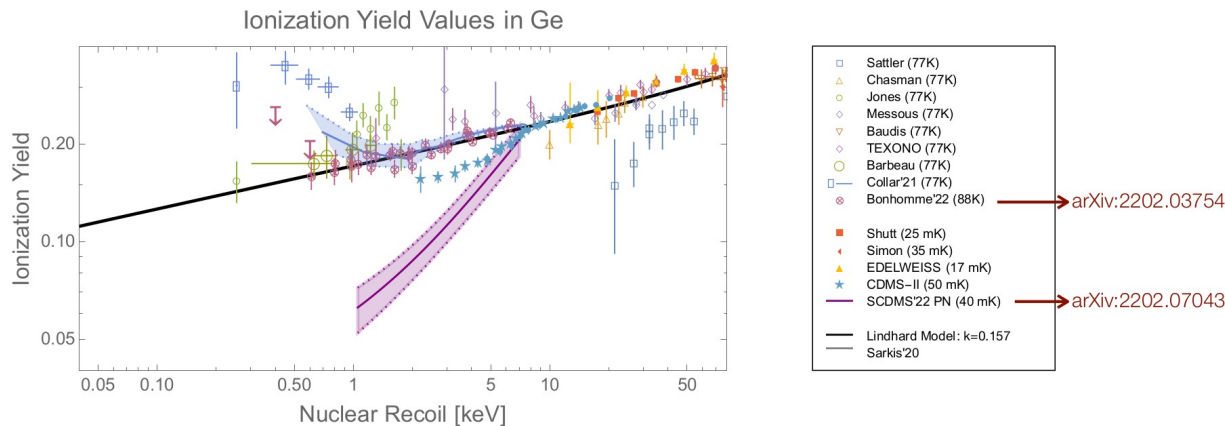


Unshielded
HPGe detector



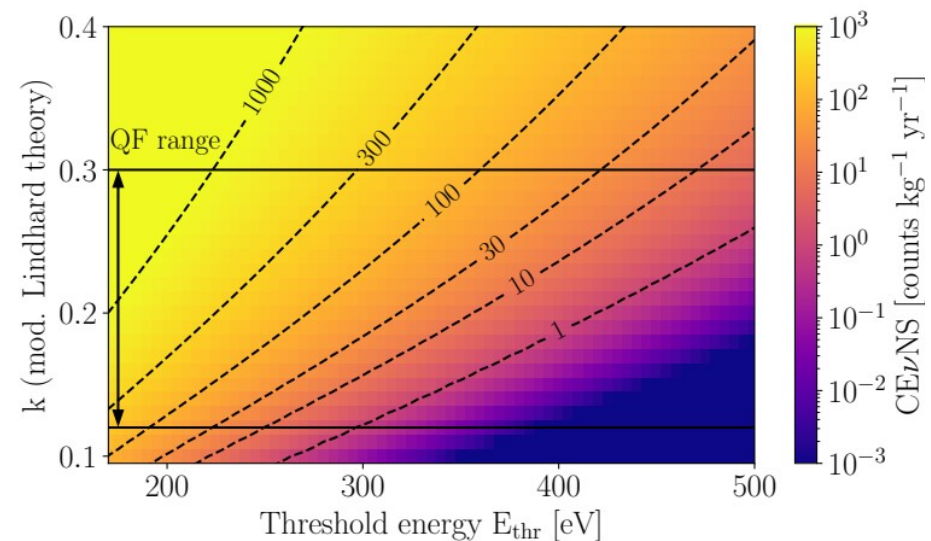
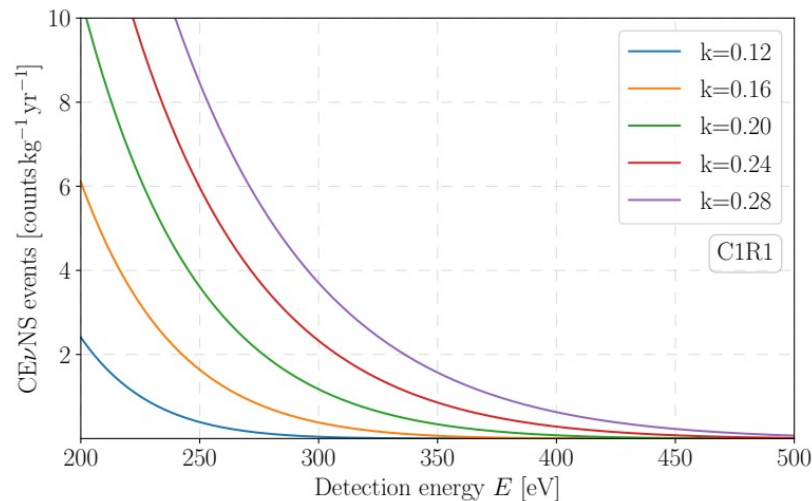
Impact of quenching at low energy

Overview of quenching measurements: T. Saab \ EXCESS 2022 \ February 16, 2022



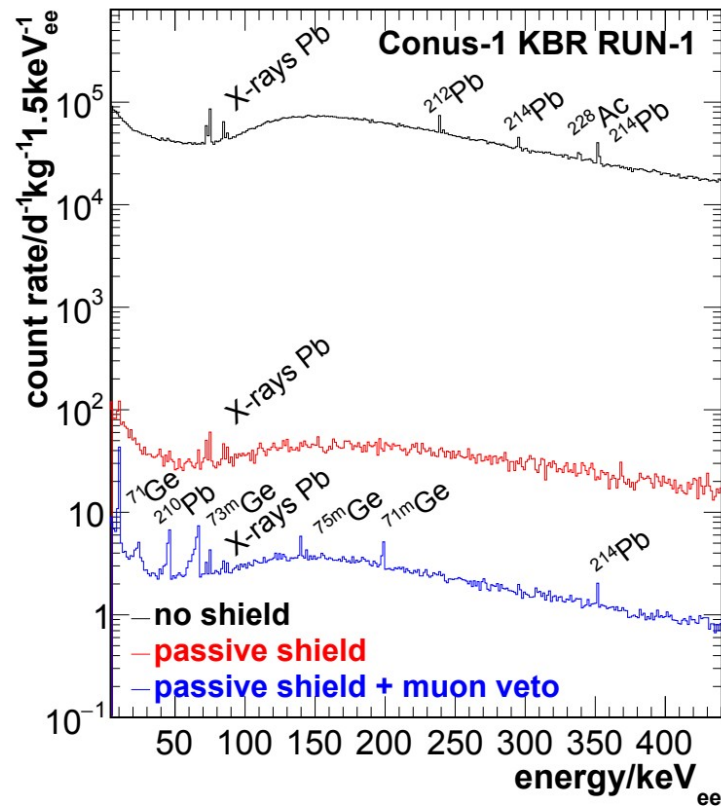
Affects sensitivity of CEvNS and DM searches:
(Collar '21) used in [arXiv:2202.09672](https://arxiv.org/abs/2202.09672)
"Suggestive evidence for Coherent Elastic Neutrino-Nucleus Scattering from reactor antineutrinos"

Quenching according to mod. Lindhard model:



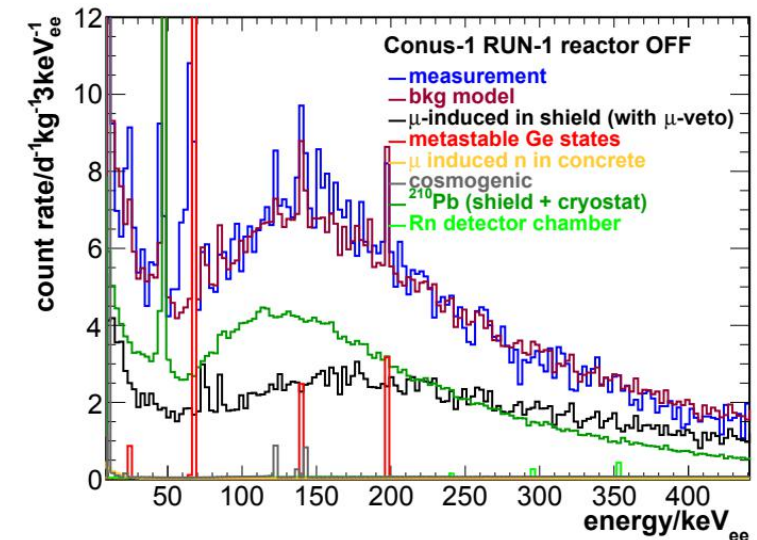
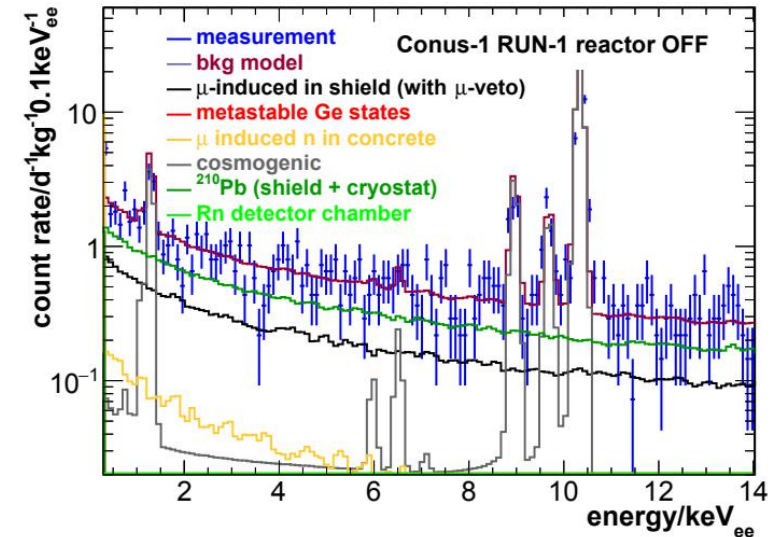
[TR, PhD thesis, 2022]

Background (model) at reactor site



Fast neutron classes	Corr. with therm. power
μ -ind. in Pb inside shield	No
μ -ind. above ceiling	No
(α, n) -reactions from walls	No
fission n from spent fuel rods	No
fission n from reactor core	Yes

[Bonet et al., arXiv:2112.09585, 2021]



Data analyses of experimental data

(Binned log)-Likelihood in the SM case:

- ON and OFF data are fitted together:

$$\log \mathcal{L} = \log \mathcal{L}_{\text{ON}} + \log \mathcal{L}_{\text{OFF}} + \text{pull terms}$$

with $\log \mathcal{L}_{\text{ON}}(s, b, \Theta_{\text{thr}_1}, \Theta_{\text{thr}_2}, \Theta_{\text{rea}}, \Theta_{\text{det}}, \Theta_{\Delta E})$
 $\log \mathcal{L}_{\text{OFF}}(b, \Theta_{\text{thr}_1}, \Theta_{\text{thr}_2}, \Theta_{\text{det}}, \Theta_{\Delta E})$

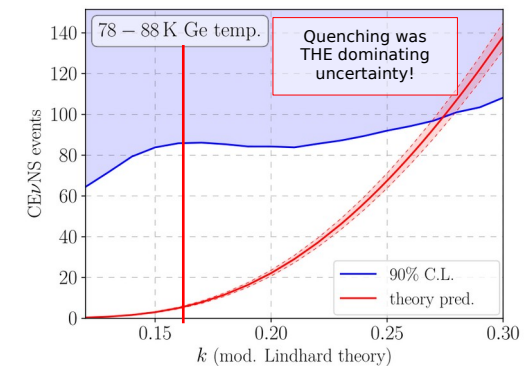
- Parameter list:
 - s : signal strength (free), b : MC normalisation (free)
 - $\Theta_{\text{thr}_1/2}$: electronic noise (free)
 - Θ_{rea} : reactor neutrino flux ($\sim 3\%$)
 - Θ_{det} : detector/DAQ (1-5%), $\Theta_{\Delta E}$: energy scale (10-20 eV_{ee})

BSM investigations:

- Extend existing LH with BSM contributions
→ CEvNS now additional background
- Additional systematic due to ν -e scatterings @[2,8]keV
→ 5% shape uncertainty on background model
- BSM limits in terms of quenching factor!

Data sets: 5 detectors (in 2 runs)

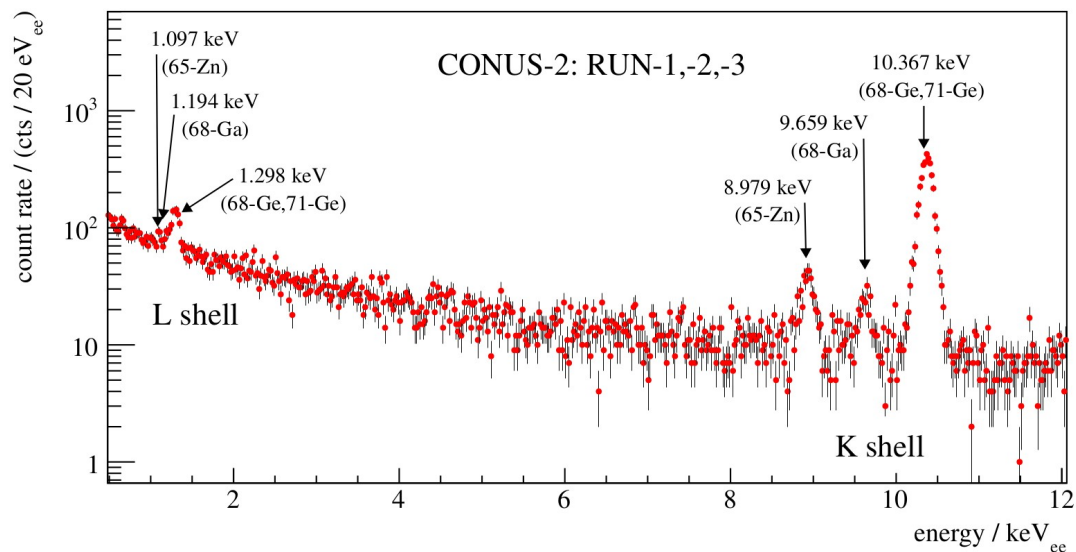
Data set	ON [kg d]	OFF [kg d]	ROI [keV]
C1R1	96.7	13.8	0.296 - 0.75
C2R1	14.6	13.4	0.311 - 1.00
C3R1	97.5	10.4	0.333 - 1.00
C1R2	19.6	12.1	0.348 - 0.75
C3R2	20.2	9.1	0.343 - 1.00
All	248.7	58.8	



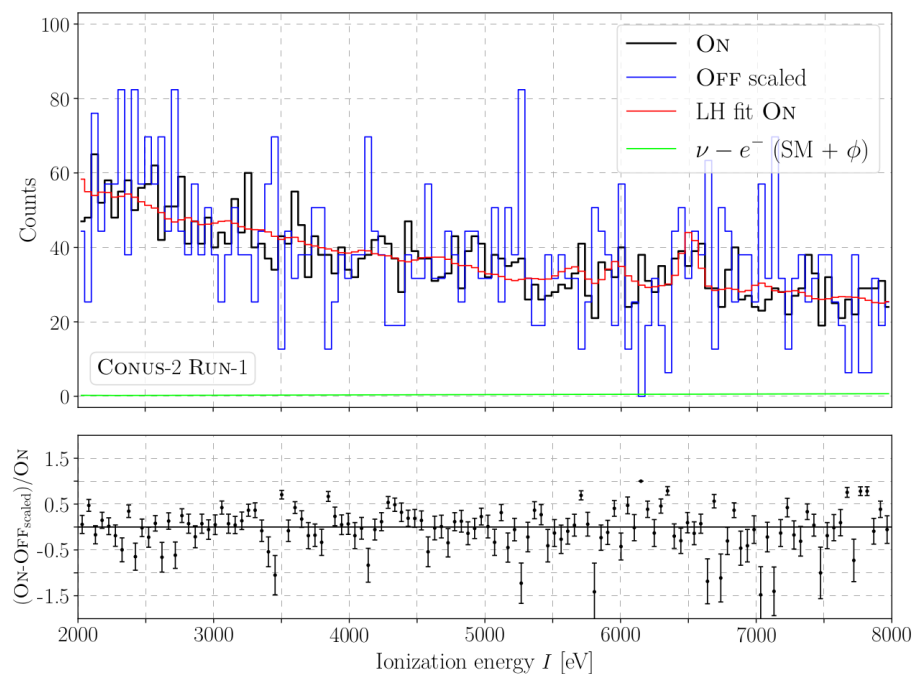
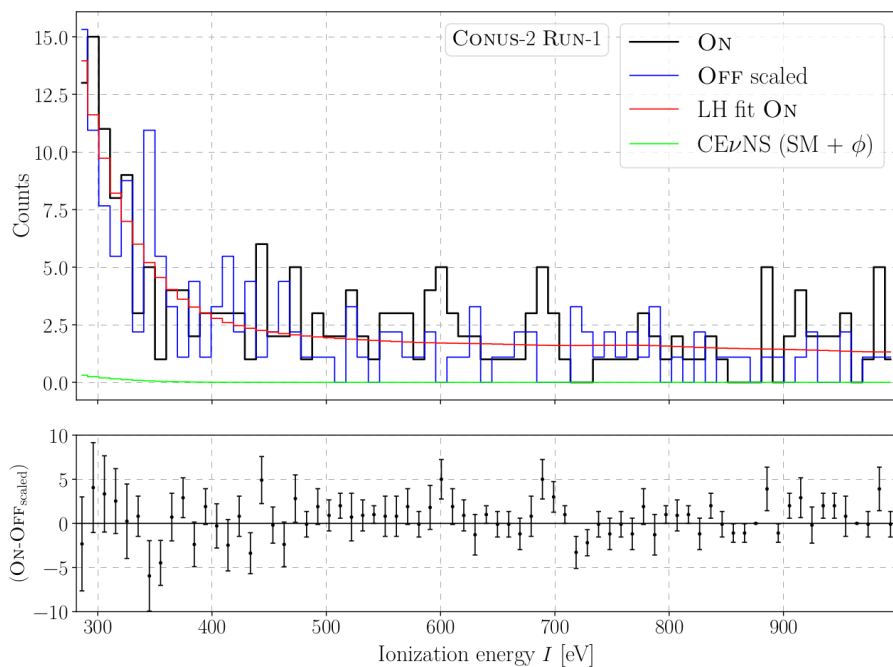
Analysis	Data set	ON [kg d]	OFF [kg d]	ROI [keV]
BSM ($\bar{\nu}_e + A(Z, N)$)	C1R1	96.7	13.8	0.276 - 0.741
	C2R1	14.6	13.4	0.281 - 0.999
	C3R1	97.5	10.4	0.333 - 0.991
	all	208.8	37.6	
BSM + ν MM ($\bar{\nu}_e + e$)	C1R1	215.4	29.6	2.013 - 7.968
	C2R1	184.6	32.2	2.006 - 7.990
	C3R1	248.5	31.7	2.035 - 7.989
	all	648.5	93.5	
ν MM ($\bar{\nu}_e + e$)	C1R2	19.6	18.5	2.010 - 7.955
	C3R2	20.8	19.0	2.007 - 7.991
	all	688.9	131.5	

Exemplary spectra: CONUS-2 RUN-1

[Bonet et al., [10.1140/epjc/s10052-021-09038-3](https://arxiv.org/abs/10.1140/epjc/s10052-021-09038-3), 2021]



[Bonet et al., [arXiv:2110.02174](https://arxiv.org/abs/2110.02174), 2021]



Sensitivity of next and future data

Estimate for recent data collection period

- Bkg stability → exposure 2.7 kg*yr
- Quenching factor known! → CONUS measurement
- New DAQ → lower threshold
- PSD → background reduction of ~20%

[TR, PhD thesis, 2022]

E_{thr} [eV]	CONUS [counts]	Z	V
300	30	0.15	42
275	49	0.25	16
250	99	0.50	4.1
225	178	0.88	1.3
200	286	1.40	0.5

→ naive counting estimate: no shape, no systematics

$$\begin{aligned} s: \text{signal events} & \quad Z = s/\sqrt{b} \\ b: \text{background events} & \quad V = (s + b)/s^2 \end{aligned}$$

Theoretical CONUS upgrade

- New experimental site similar to KBR
- Same bkg and detector specifications
- Exposure: 500 kg*yr

→ precision neutrino physics with CevNS!

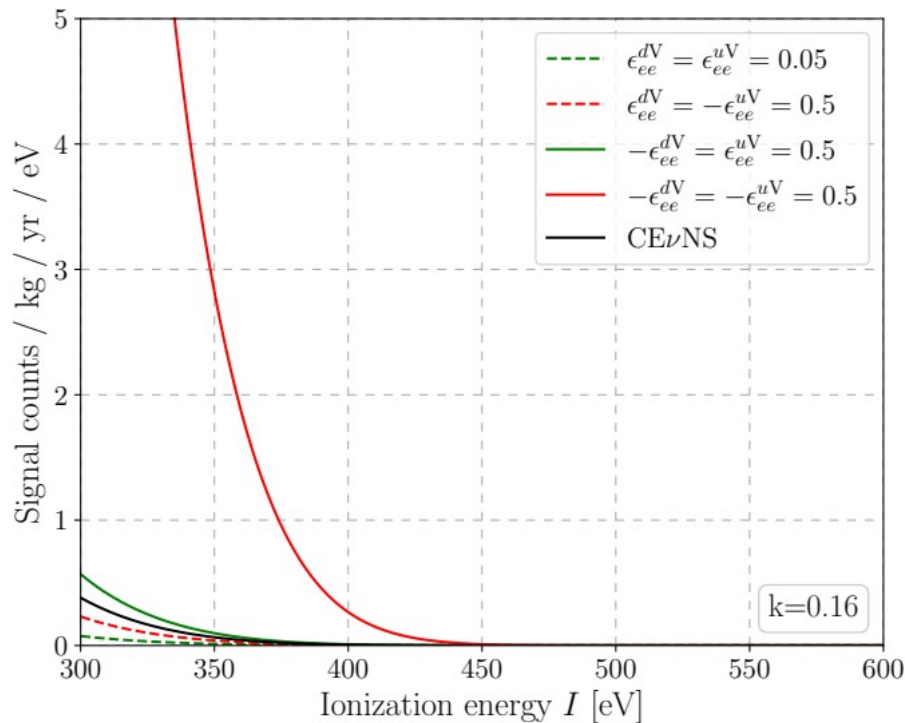
E_{thr} [eV]	CONUS-100 [counts]	Z	V
300	$5.5 \cdot 10^3$	2.1	$227 \cdot 10^{-3}$
275	$9.1 \cdot 10^3$	3.4	$85 \cdot 10^{-3}$
250	$18 \cdot 10^3$	6.8	$22 \cdot 10^{-3}$
225	$33 \cdot 10^3$	12.0	$7 \cdot 10^{-3}$
200	$53 \cdot 10^3$	19.0	$3 \cdot 10^{-3}$

Vector-type NSIs

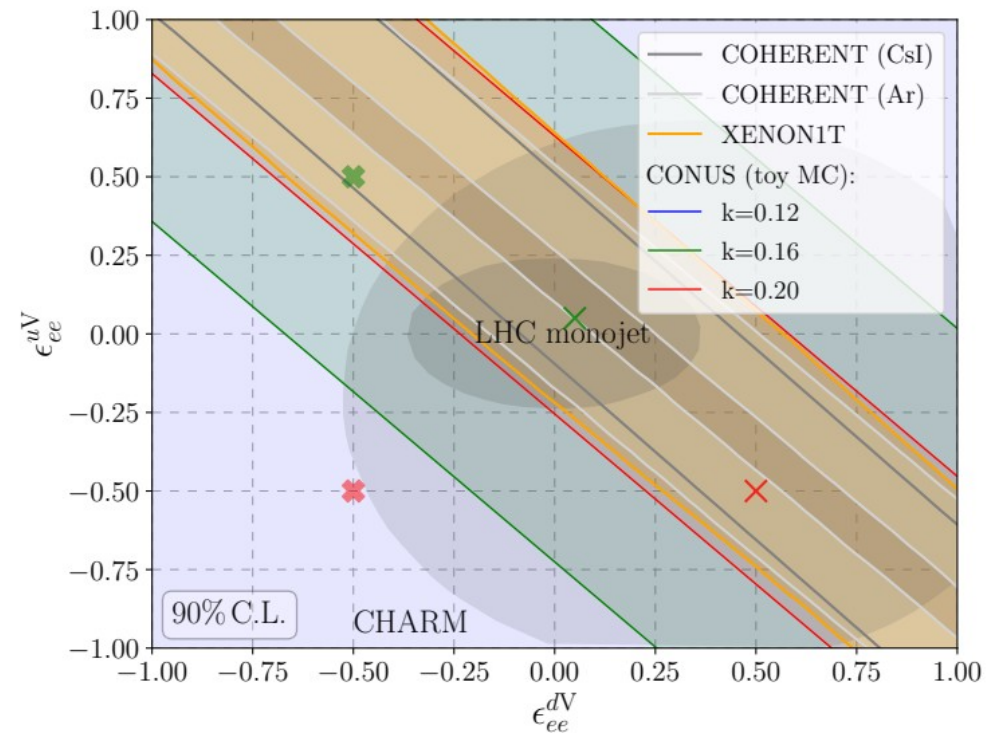
$$\mathcal{O}_{\alpha\beta}^{qV} = (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta) (\bar{q} \gamma_\mu P_{L/R} q) + \text{h.c.}$$

$$\begin{aligned} \mathcal{Q}_{NSI}^V = & (2\epsilon_{\alpha\alpha}^{uV} + \epsilon_{\alpha\alpha}^{dV} + g_p^V) Z + (\epsilon_{\alpha\alpha}^{uV} + 2\epsilon_{\alpha\alpha}^{dV} + g_n^V) N \\ & + \sum_{\alpha \neq \beta} \left[(2\epsilon_{\alpha\beta}^{uV} + \epsilon_{\alpha\beta}^{dV}) Z + (\epsilon_{\alpha\beta}^{uV} + 2\epsilon_{\alpha\beta}^{dV}) N \right] \end{aligned}$$

Expectation for C1R1



Limits of combined data sets

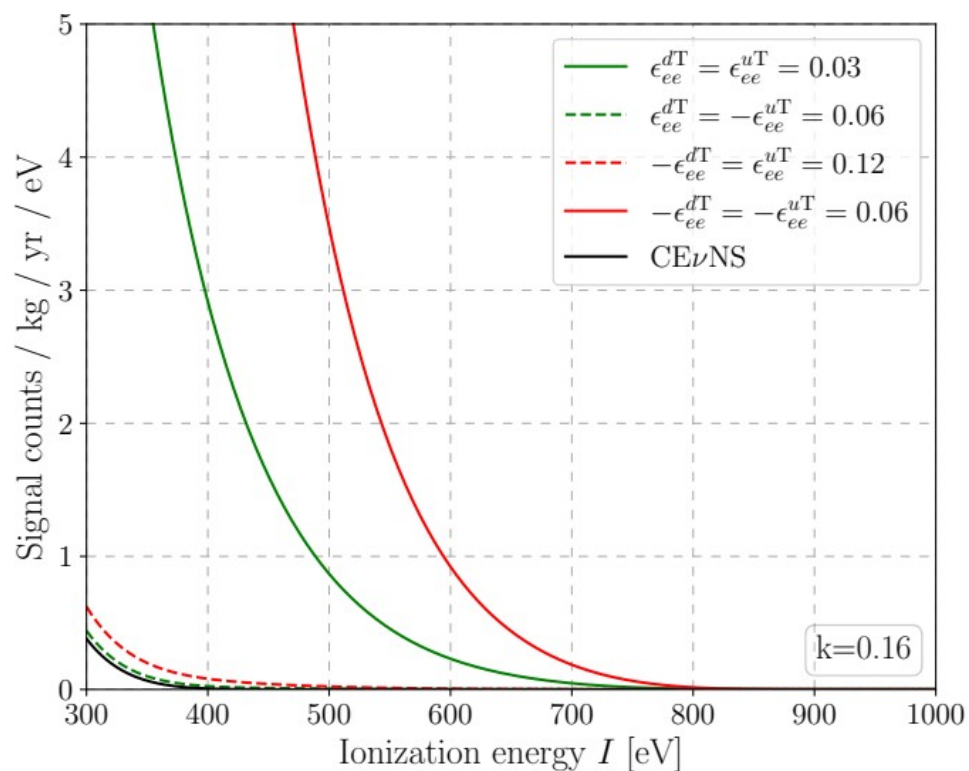


[Bonet et al., [arXiv:2110.02174](https://arxiv.org/abs/2110.02174), 2021]

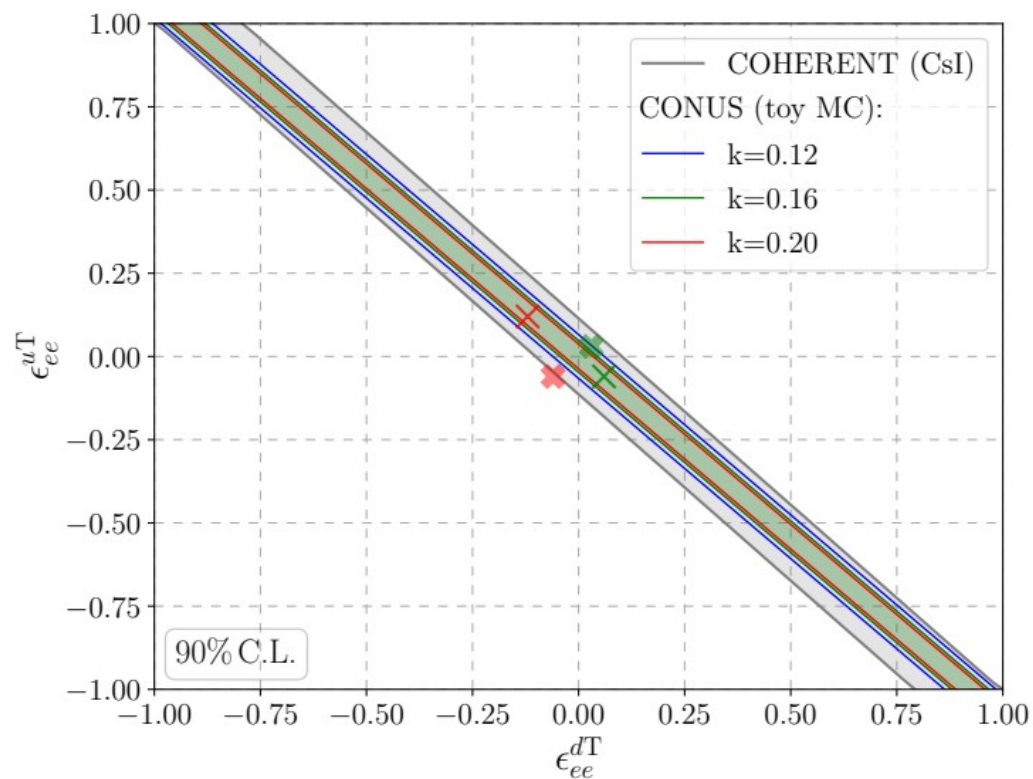
Tensor-type NSIs

$$\mathcal{O}_{\alpha\beta}^{qT} = (\bar{\nu}_\alpha \sigma^{\mu\nu} \nu_\beta) (\bar{q} \sigma_{\mu\nu} q) + \text{h.c.}$$

Expectation for C1R1



Limits of combined data sets

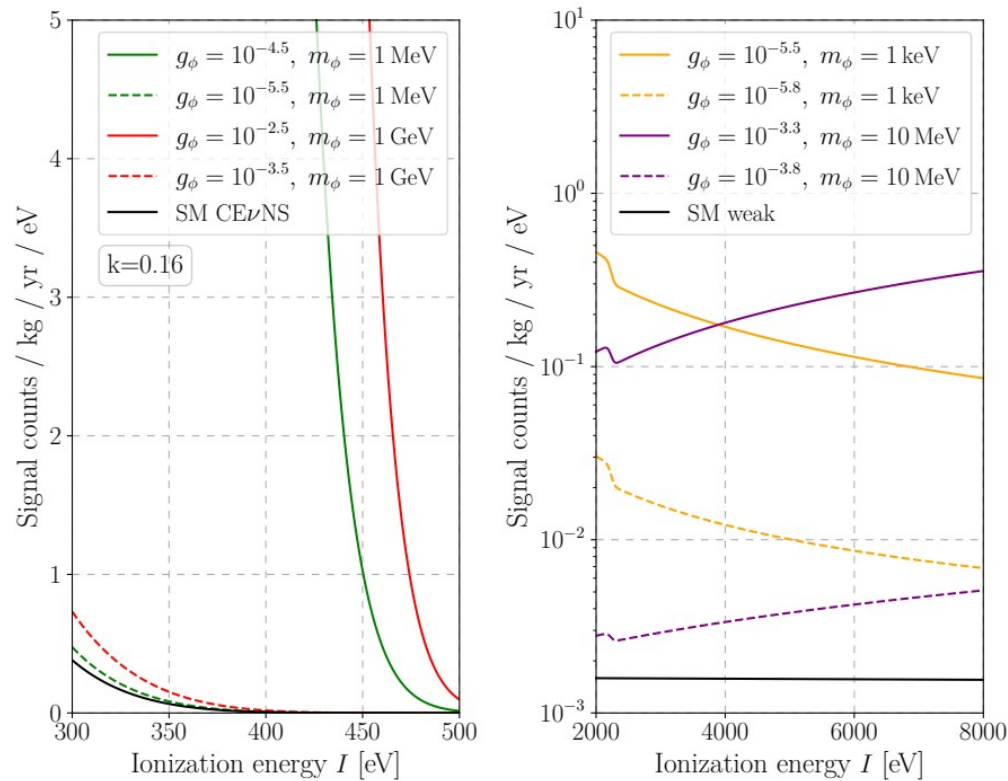


[Bonet et al., [arXiv:2110.02174](https://arxiv.org/abs/2110.02174), 2021]

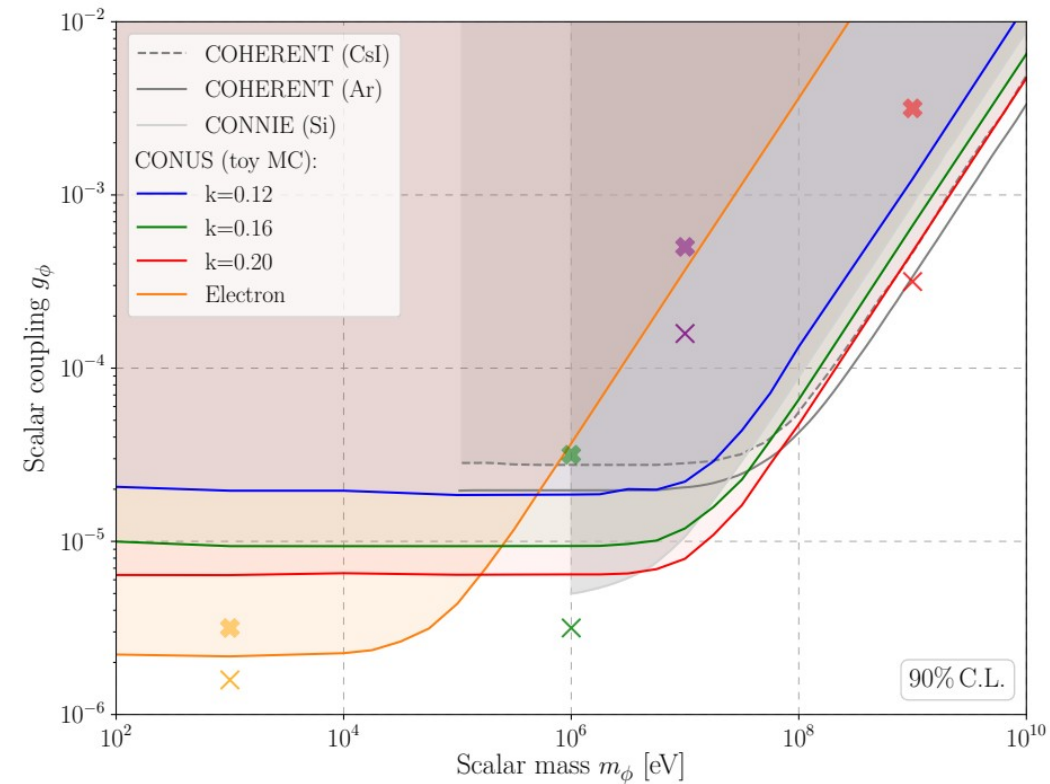
Light scalar mediator

$$\mathcal{L}_\phi = \phi \left(g_\phi^{qS} \bar{q}q + g_\phi^{eS} \bar{e}e + g_\phi^{\nu S} \bar{\nu}_R \nu_L + \text{h.c.} \right) - \frac{1}{2} m_\phi^2 \phi^2$$

Expectation for C1R1



Limits of combined data sets

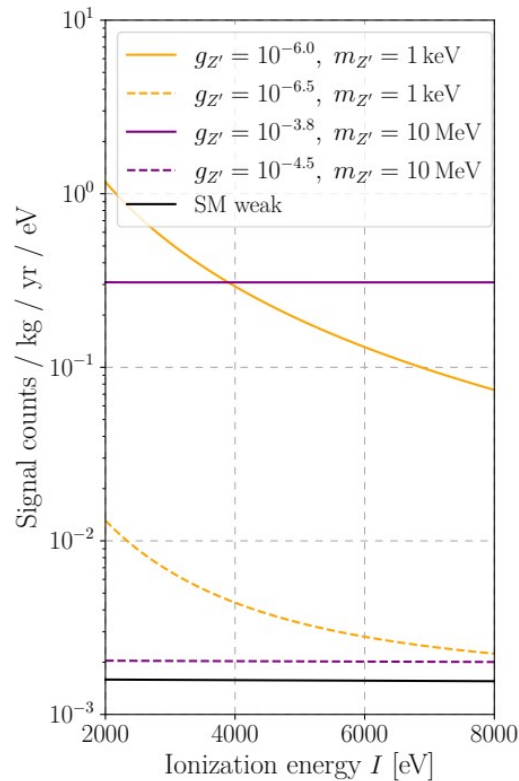
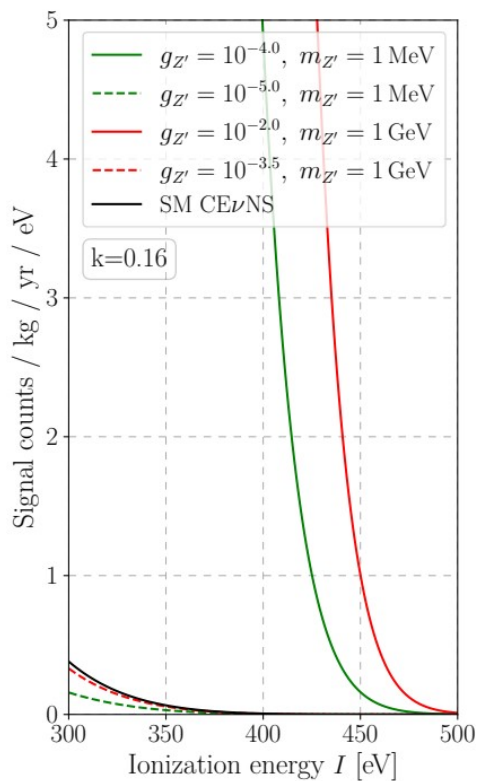


[Bonet et al., [arXiv:2110.02174](https://arxiv.org/abs/2110.02174), 2021]

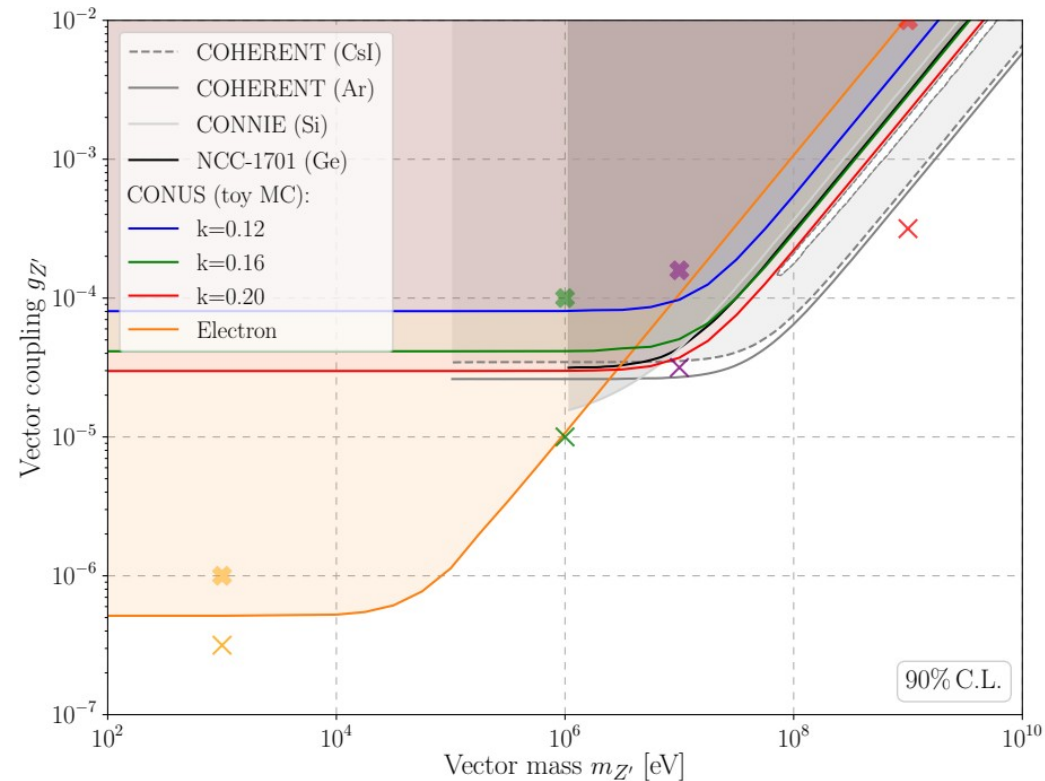
Light vector mediator

$$\mathcal{L}_{Z'} = Z'_\mu \left(g_{Z'}^{\nu V} \bar{\nu}_L \gamma^\mu \nu_L + g_{Z'}^{e V} \bar{e} \gamma^\mu e + g_{Z'}^{q V} \bar{q} \gamma^\mu q \right) + \frac{1}{2} m_{Z'}^2 Z'_\mu Z'^\mu$$

Expectation for C1R1



Limits of combined data sets



[Bonet et al., [arXiv:2110.02174](https://arxiv.org/abs/2110.02174), 2021]